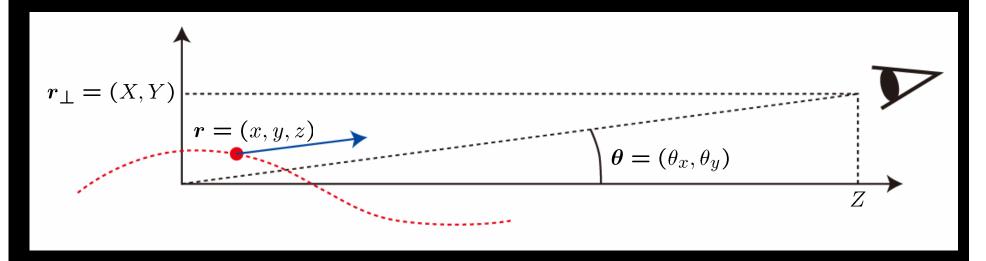
## Light Source II

Takashi TANAKA (RIKEN SPring-8 Center)

# Characteristics of SR (2) - Radiation from IDs

## Electron Trajectory in ID Field

## Coordinate Systems



SR emitted by an electron moving at  $\mathbf{r} = (x,y,z)$ Observation of SR at  $\mathbf{R} = (X,Y,Z)$ 

If the far-field approximation (|r| < Z) is applicable, the radiation pattern depends only on the observation angle  $\theta = (\theta_x, \theta_y)$ .

### Field Integrals

$$\frac{d\mathbf{P}}{dt} = m\gamma \frac{d\mathbf{v}}{dt} = -e\mathbf{v} \times \mathbf{B} \longrightarrow \begin{cases} m\gamma \dot{v_x} = -e(v_y B_z - v_z B_y) \\ m\gamma \dot{v_y} = -e(v_z B_x + v_x B_z) \end{cases}$$

Equation of motion of an electron moving in a magnetic field **B** 

$$B_z \equiv 0$$

$$m\gamma \frac{dv_{x,y}}{v_z dt} = m\gamma \frac{dv_{x,y}}{dz} = \pm eB_{y,x}$$

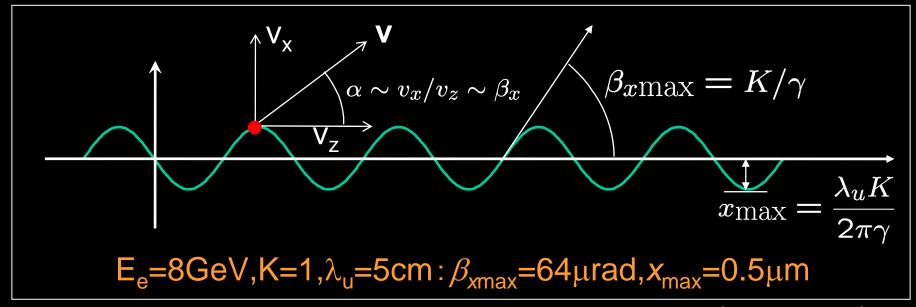
$$\beta_{x,y} = \pm \frac{e}{\gamma mc} \int^{z} B_{y,x}(z') dz' \equiv \pm \frac{e}{\gamma mc} I_{1y,1x}(z)$$
$$x, y = \pm \frac{e}{\gamma mc} \int^{z} \int^{z'} B_{y,x}(z'') dz'' \equiv \pm \frac{e}{\gamma mc} I_{2y,2x}(z)$$

 $I_1, I_2$ : 1st and 2nd field integrals of ID

## Electron Trajectory in an Ideal ID

$$\begin{cases} B_x(z) = 0 \\ B_y(z) \sim B_0 \sin\left(\frac{2\pi z}{\lambda_u}\right) \end{cases} \begin{cases} \beta_y = 0 \\ \beta_x = \frac{K}{\gamma} \cos\left(\frac{2\pi z}{\lambda_u}\right) \end{cases} \begin{cases} y = 0 \\ x = \frac{\lambda_u K}{2\pi \gamma} \sin\left(\frac{2\pi z}{\lambda_u}\right) \end{cases}$$
 magnetic field velocity position

$$K=rac{eB_0\lambda_u}{2\pi mc}$$
 K value, Deflection parameter



#### Effects due to the ID Field

transverse velocity  $\beta_x = \frac{K}{\gamma} \cos\left(\frac{2\pi z}{\lambda_u}\right)$  longitudinal velocity  $\beta_z = \sqrt{\beta^2 - \beta_x^2} \quad \text{total velocity}$   $= 1 - \frac{1}{2\gamma^2} - \frac{K^2}{4\gamma^2} - \frac{K^2}{4\gamma^2} \cos\left(\frac{4\pi z}{\lambda_u}\right)$   $\bar{\beta}_z : \text{average velocity} \quad \text{oscillating component}$ 

#### ID field induces:

- transverse(x) oscillation
- longitudinal (z) oscillation
- effective deceleration( $\Delta \beta_z = K^2/4\gamma^2$ )

#### **Electron Motion: Two Forms**

$$\beta_x = \frac{K}{\gamma} \cos\left(\frac{2\pi z}{\lambda_u}\right)$$

- Horizontal oscillation with a period of  $\lambda_u$
- Major contribution to radiation

$$\beta_z = \bar{\beta_z} - \frac{K^2}{4\gamma^2} \cos\left(\frac{4\pi z}{\lambda_u}\right)$$

- Longitudinal oscillation with a period of  $\lambda_u/2$
- Amplitude  $1/\gamma$  times lower than  $\beta_x$ .
- Minor contribution, but source of vertical polarization observed vertically off-axis.

## General Form of Time Squeezing

$$\frac{d\tau}{dt} = 1 - \beta \cdot \mathbf{n}$$

$$\beta_z = \sqrt{\beta^2 - \beta_x^2 - \beta_y^2}$$

$$\sim 1 - (\gamma^{-2} + \beta_x^2 + \beta_y^2)/2$$

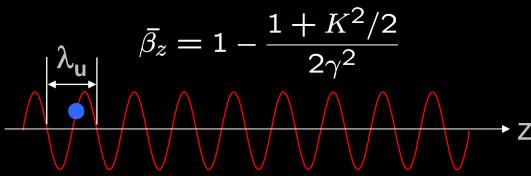
$$n_z \sim 1 - (\theta_x^2 + \theta_y^2)/2$$

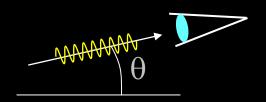
$$= \frac{1}{2\gamma^2} + (\theta_x - \beta_x)^2 + (\theta_y - \beta_y)^2$$

Time squeezing takes place most significantly when the direction of the electron motion coincides with that of observation ( $\beta = \theta$ ).

## Qualitative Descriptions of Undulator Radiation

## Fundamental Wavelength





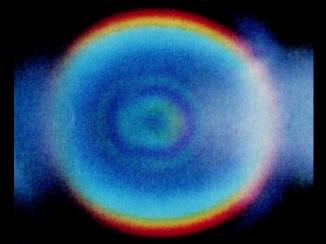
$$T = \lambda_u/v_z = \lambda_u/c$$



period of electron motion time squeezing
= period of emitted light

$$T' = T(1 - \bar{\beta}_z \cos \theta)$$
 period of observed light



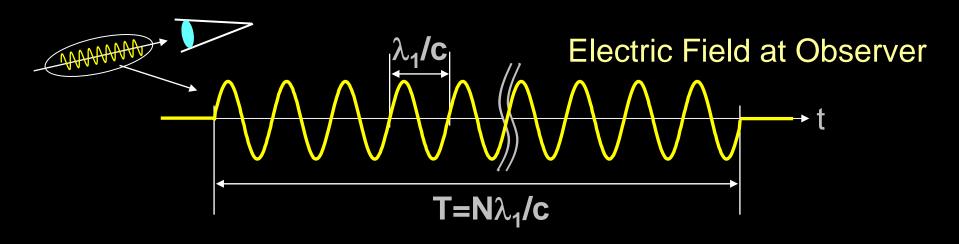


Fundamental Wavelength  $\lambda_1$ 

$$\lambda_1 = \lambda_u (1 - \bar{\beta}_z \cos \theta)$$
$$= \frac{\lambda_u}{2\gamma^2} (1 + \gamma^2 \theta^2 + K^2/2)$$

H. Kitamura et al.,J. Appl. Phys. 21 (1982) 1728

#### Effects due to Finite Periods



$$E(t) = \begin{cases} E_0 \sin \omega_1 t & ; -T/2 \le t \le T/2 \\ 0 & ; t < -T/2, T/2 < t \end{cases}, \ \omega_1 = 2\pi c/\lambda_1$$

**Fourier Transform** 

$$\frac{d^2 F(\omega, \theta)}{d\Omega d\omega/\omega} \propto |\tilde{E}(\theta, \omega)|^2 \propto \text{sinc}^2 \left[ \pi N \frac{\omega - \omega_1(\theta)}{\omega_1(\theta)} \right]$$

Square of "sinc" function dominates the UR

### Brief Note on UR Formulae

- In the previous derivations of UR spectral function, no knowledge on electrodynamics is required.
- In practice,  $E_0$  is a complicated function of  $\theta$  and K, and needs to be calculated by Fourier transforming the electric field derived from the Lienard-Wiecherd potential.
- However, the simple derivation gives us a clear understanding on UR properties.

## Energy and Angular Profile of UR

$$\frac{d^2 F(\omega, \theta)}{d\Omega d\omega/\omega} = F_0 \text{sinc}^2 \left[ \pi N \frac{\omega - \omega_1(\theta)}{\omega_1(\theta)} \right]$$

#### Energy Profile at $\theta = 0$

$$F_0 \mathrm{sinc}^2(N\pi\varepsilon)$$
  
;  $\varepsilon = [\omega - \omega_1(0)]/\omega_1(0)$ 

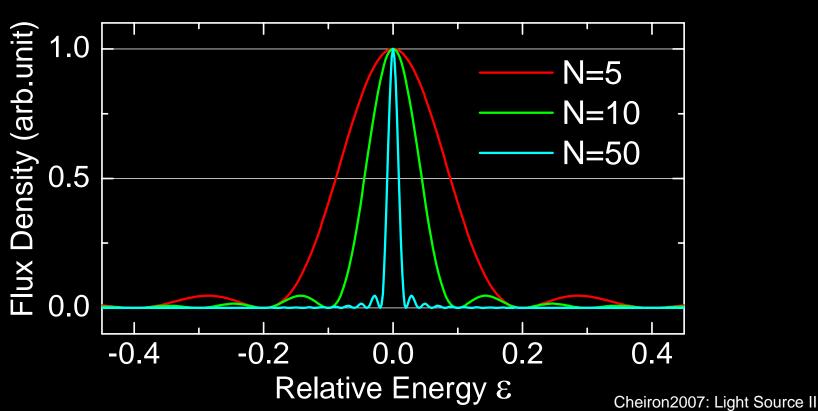
#### Angular Profile at $\omega = \alpha \omega_1(0)$

$$F_0 \operatorname{sinc}^2[N\pi(\alpha\Theta^2 + \alpha - 1)]$$
  
;  $\Theta = \gamma\theta/\sqrt{1 + K^2/2}$ 

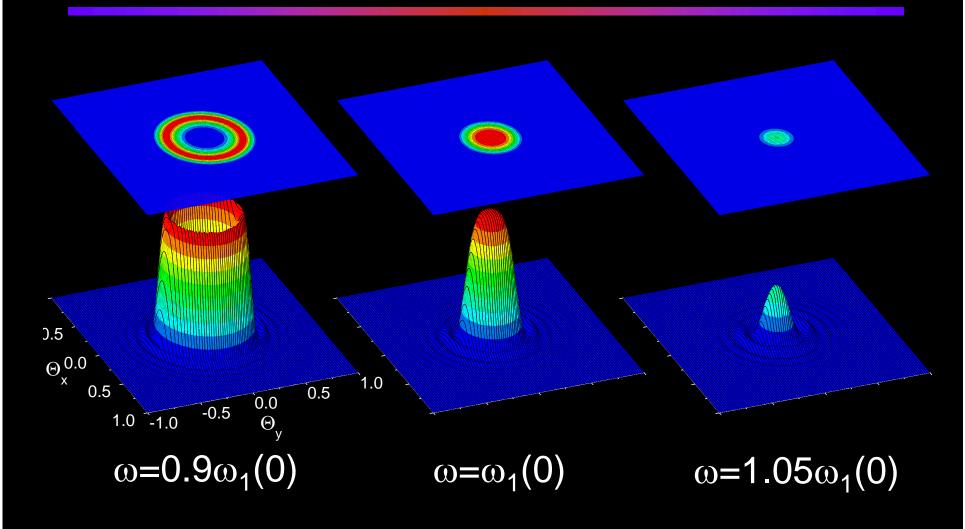
## **Energy Profile: Example**

$$\frac{d^2 F(\omega, \theta)}{d\Omega d\omega/\omega} = F_0 \text{sinc}^2(N\pi\varepsilon); \quad \text{sinc}^2(2.783) \sim 1/2$$





## Angular Profile: Example



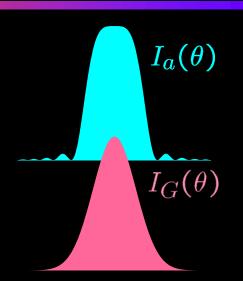
## Angular Divergence and Beam Size

Angular Profile at  $\omega = \omega_1(0)$ 

$$I_a(\theta) = F_0 \operatorname{sinc}^2 \left[ \frac{\pi N(\gamma \theta)^2}{1 + K^2/2} \right]$$

approximation

Gaussian Profile with  $\sigma_{r'}$   $I_G(\theta) = F_0 \exp(-\theta^2/2\sigma_{r'}^2)$ 



$$\sigma_{r'} = \sqrt{rac{1+K^2/2}{4N\gamma^2}} = \sqrt{rac{\lambda_1}{2L}}$$
 Angular Divergence of UR

**Diffraction Limit** 

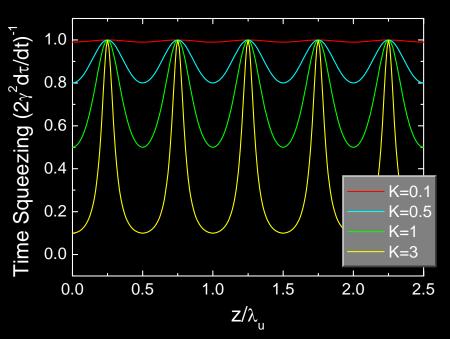
$$\sigma_r = rac{\lambda_1}{4\pi\sigma_{rl}} = rac{\sqrt{\lambda_1 L}}{4\pi}$$
 Beam Size of UR

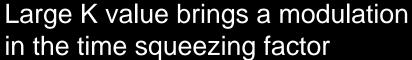
## Higher Harmonics

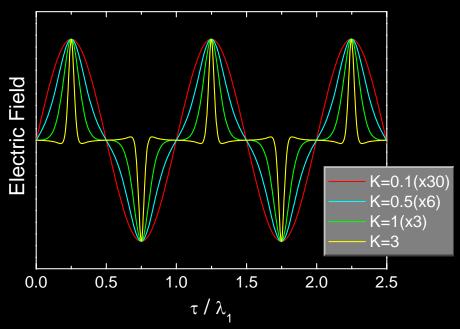
- In addition to the fundamental radiation at  $\omega_1$ , higher-energy radiation at  $n\omega_1$ , called higher harmonics, is observed. The integer n is referred to as a harmonic number.
- This is a consequence of the fact that the time-squeezing factor depends on the longitudinal electron position and thus the electric field in the time domain is distorted.

## Interpretation of Higher Harmonics

$$\frac{d\tau}{dt} = 1 - \beta \cdot n = \frac{1}{2\gamma^2} \left[ 1 + K^2 \cos^2(2\pi z/\lambda_u) \right]$$
 on-axis observation:n=(0,0,1)



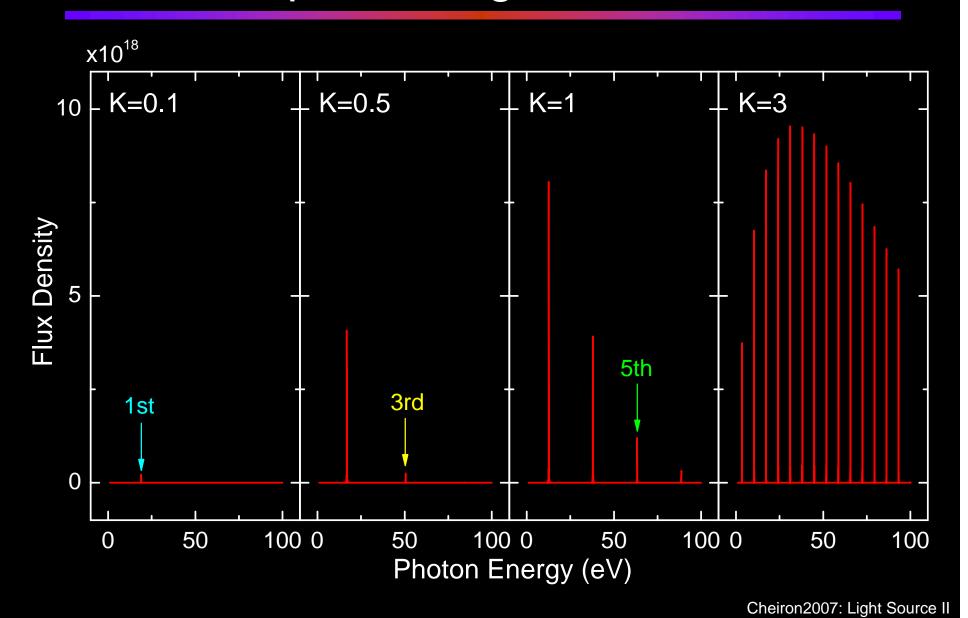




Distortions of the electric field takes place due to the nonuniform time squeezing. Due to symmetry, even harmonics do not appear.

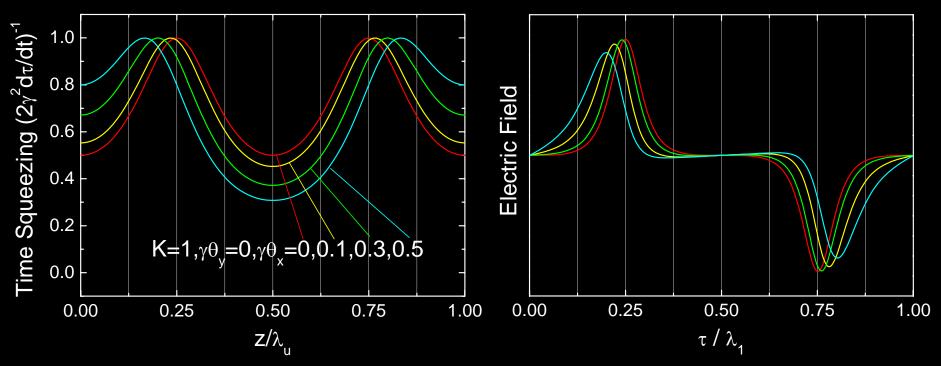
Cheiron2007: Light Source II

## Examples of Higher Harmonics



## Even Harmonics Horizontally Off Axis

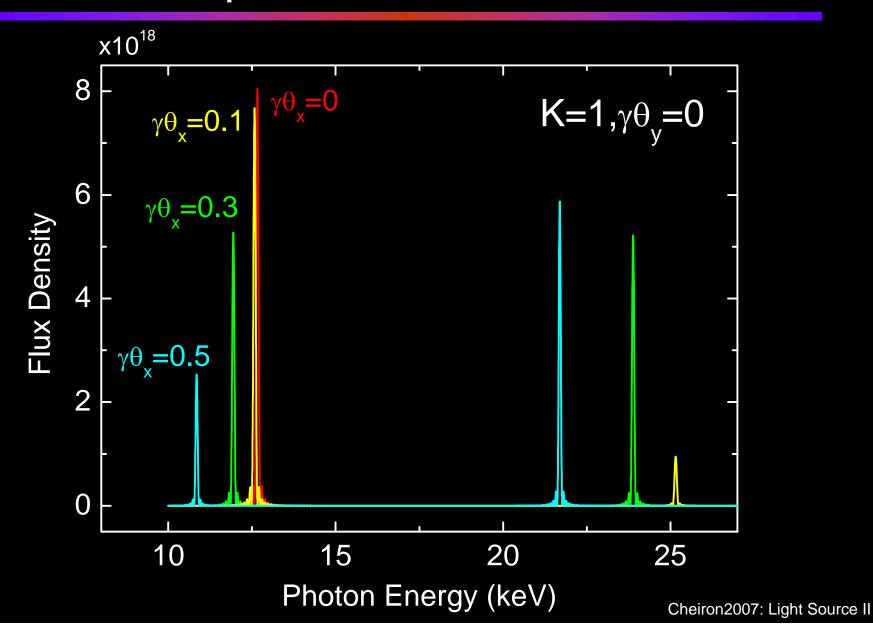
$$\frac{d\tau}{dt} = \frac{1}{2\gamma^2} \left[ 1 + \left( \gamma \theta_x - K \cos \frac{2\pi z}{\lambda_u} \right)^2 \right]$$



The position for the maximum time squeezing is shifted due to finite  $\theta_x$ .

The symmetry of the electric field is broken, resulting in appearance of even harmonics.

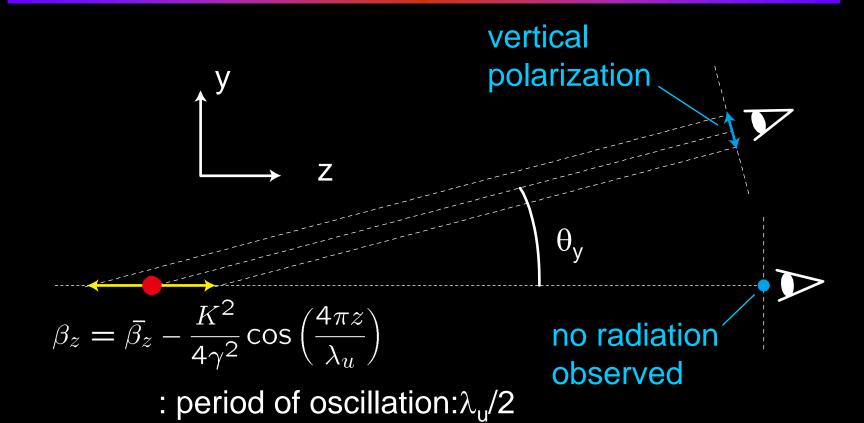
## **Examples of Even Harmonics**



## Even Harmonics Vertically Off Axis

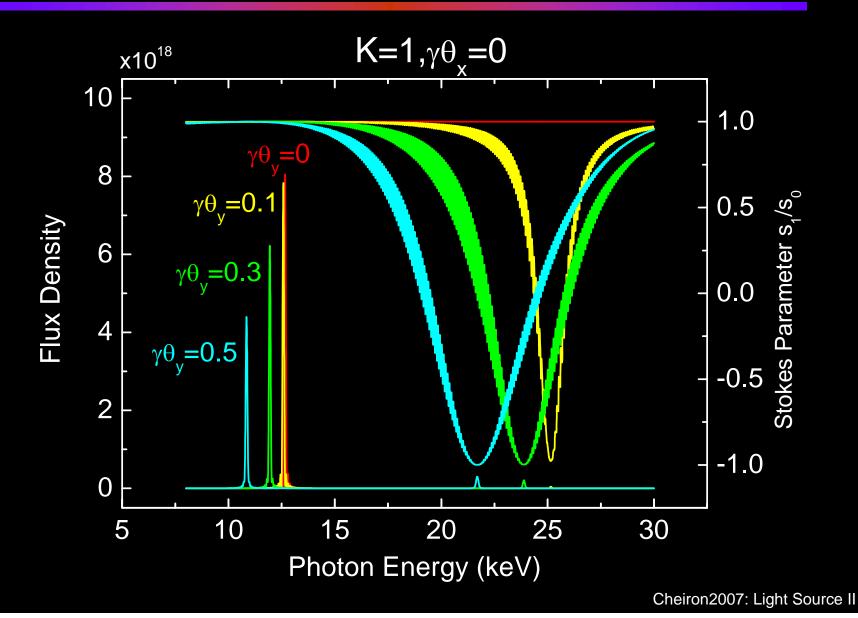
- Vertically off-axis observation does not break the symmetry of the E-field.
- Nevertheless, even harmonics are observed due to the longitudinal oscillation in electron motion with a period of  $\lambda_u/2$ .
- Such even harmonics are vertically polarized, reflecting the electron motion projected onto the plane of observation.

### Mechanism of Vertical Polarization



Note: amplitude of oscillation is  $\sim \gamma^{-1}$  smaller than that of  $\beta_x$ 

## Example of Vertical Polarization



## Optical Properties of Higher Harmonics

For the n-th harmonic radiation,

$$\frac{d^2 F(\omega, \theta)}{d\Omega d\omega/\omega} = F_0 \text{sinc}^2 \left[ \pi n N \frac{\omega - n\omega_1(\theta)}{n\omega_1(\theta)} \right]$$



$$\left. rac{\Delta \omega}{n \omega_1(0)} \right|_{FWHM} \sim \left. rac{0.8858}{nN} 
ight.$$
 band width  $\sigma_{r'n} = \sqrt{rac{1+K^2/2}{4nN\gamma^2}} = \sqrt{rac{\lambda_1/n}{2L}}$  angular divergence  $\sigma_{rn} = rac{\lambda_1/n}{\Delta r} = \sqrt{rac{L\lambda_1/n}{2L}}$  beam size

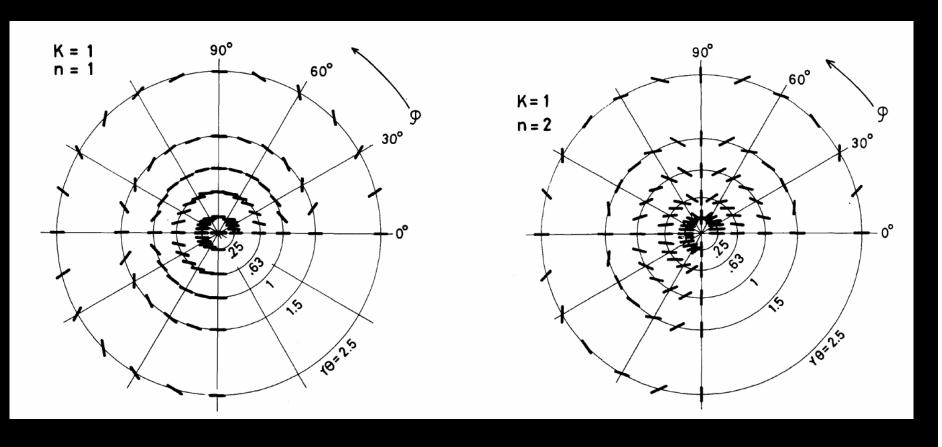
### **Polarization**

- No circular polarized radiation (CPR) is observed unlike the BM radiation.
- This is due to cancellation of CPR components between two adjacent half-period with opposite direction of electron motion (rotation).
- The direction of the linear polarization observed off axis is tilted due to the longitudinal oscillation of electron motion.

## Polarization: Examples

Examples of the direction of linear polarization for various observation angles.

H. Kitamura, JJAP 19 (1980) L185

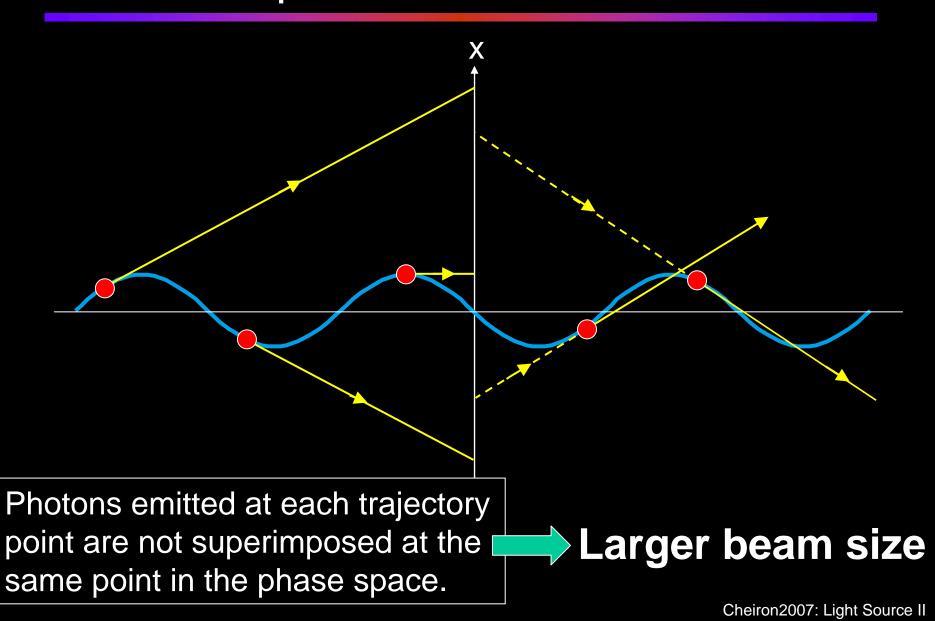


# Qualitative Descriptions of Wiggler Radiation

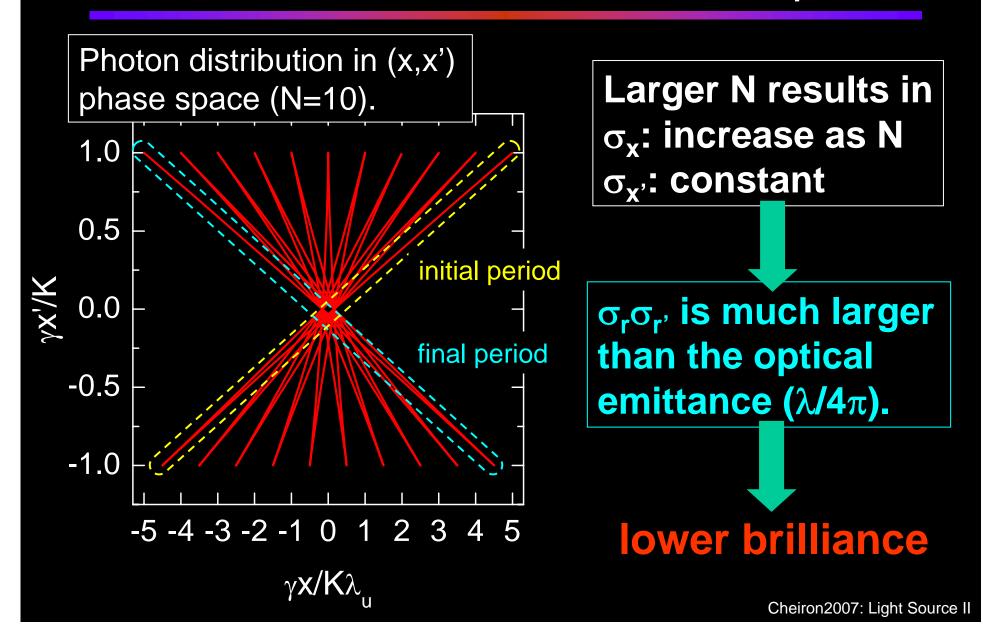
#### WR: Incoherent Sum of BM Radiation

- Wiggler radiation (WR) is regarded to be incoherent superposition of SR emitted at each position.
- Thus, the flux (density) of WR is simply 2N times that of BM radiation.
- It should be noted, however, that the brilliance of WR is much lower than is expected from a simple estimation.

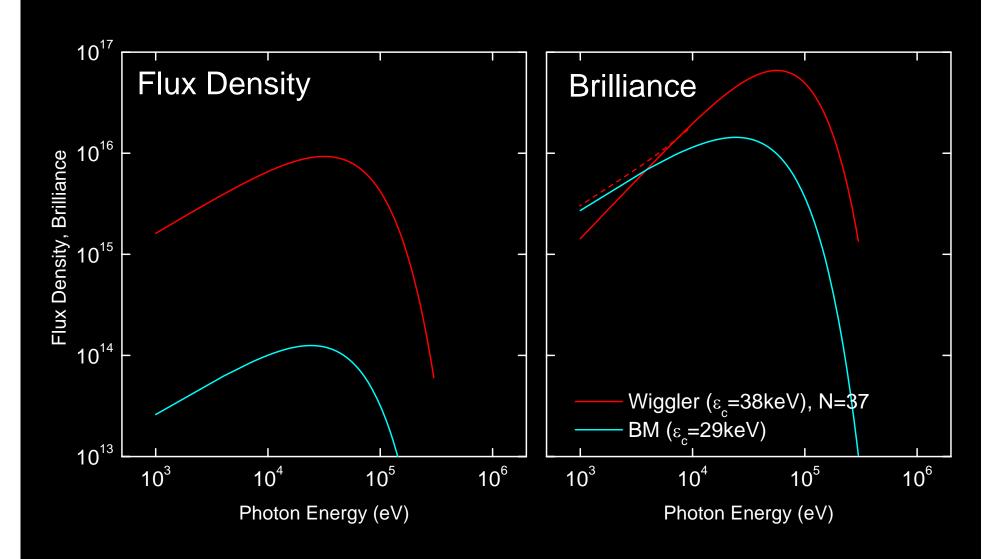
## Multiple Source Point in WR

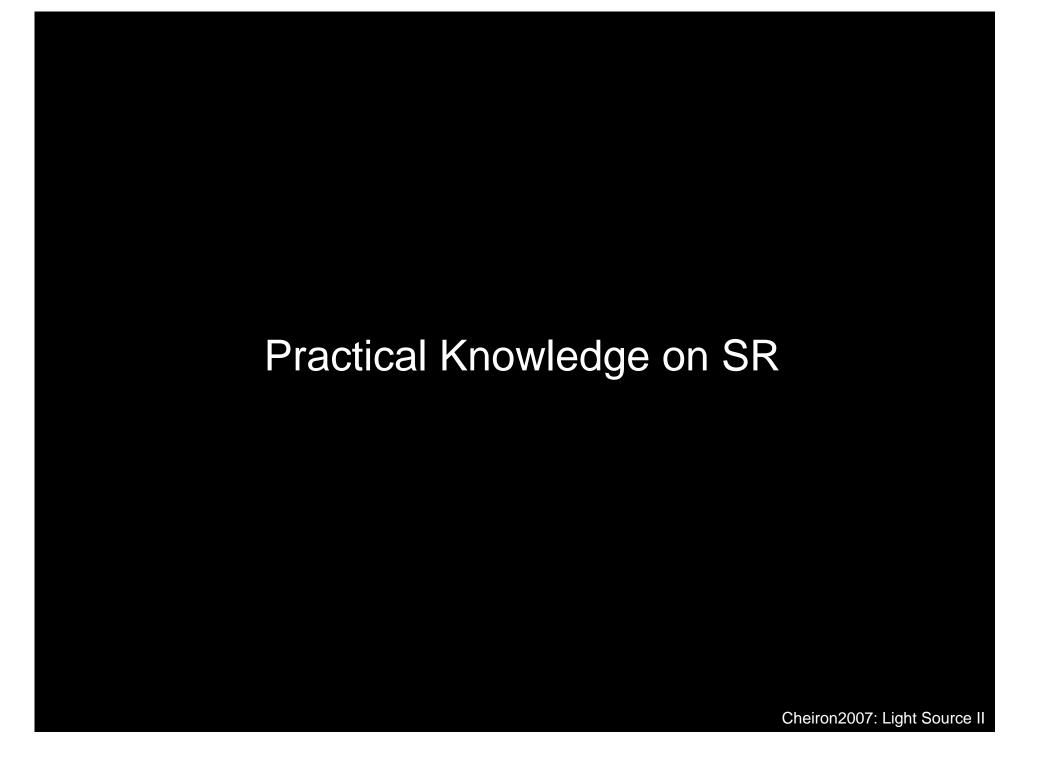


## Photon Distribution in Phase Space

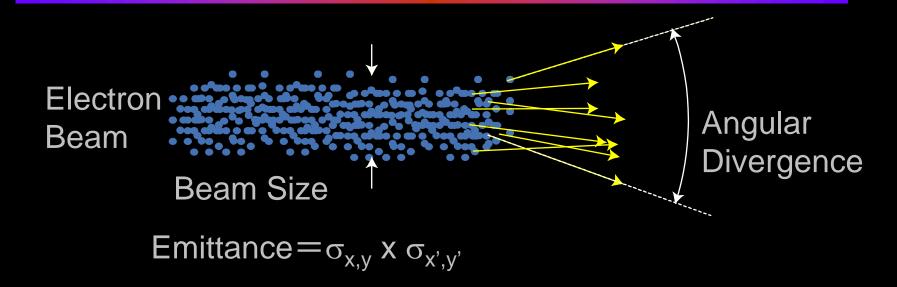


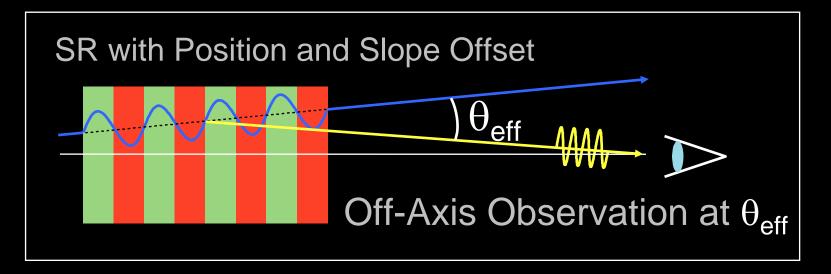
## Comparison with BM Radiation





## Effects due to Finite Emittance (1)





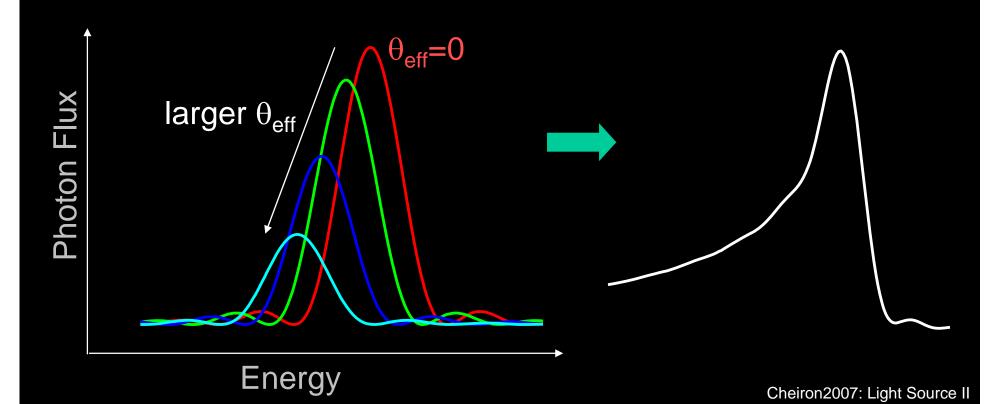
## Effects due to Finite Emittance (2)

Off-axis observation at  $\theta_{eff}$ 

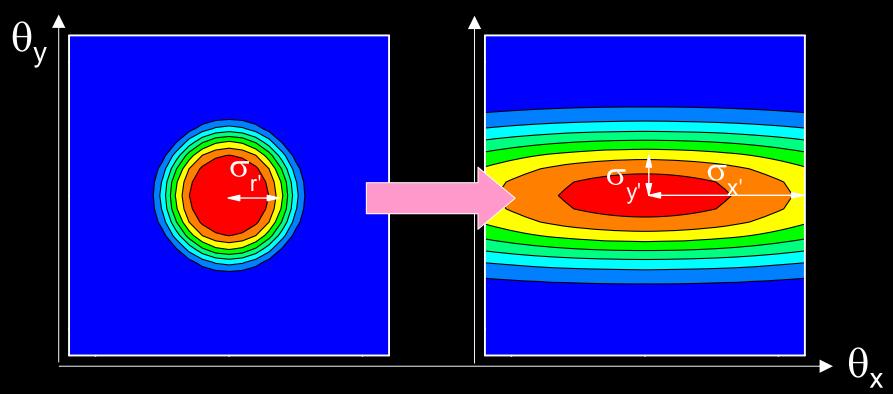


Peak shift to lower energy

$$\omega_1(\theta) = \frac{4\pi c \gamma^2 / \lambda_u}{1 + \gamma^2 \theta^2 + K^2 / 2}$$



## Effects due to Finite Emittance (3)



Under Gaussian approximation

$$\sigma_{x',y'} = \sqrt{\sigma_{r'}^2 + \sigma_{ex',ey'}^2}, \ \sigma_{x,y} = \sqrt{\sigma_r^2 + \sigma_{ex,ey}^2}$$

This expression is valid only near the resonance energy ( $n\omega_1$ ).

## Effects due to Finite Emittance (4)

Simple scheme to estimate the on-axis flux density and brilliance.

$$F = F_0 \times 2\pi\sigma_{r'}^2$$

F<sub>0</sub>: on-axis flux density for zero-emittance beam F: total flux integrated over whole solid angle

$$F_e = \frac{F}{2\pi\sigma_{x'}\sigma_{y'}} = F_0 \frac{\sigma_{r'}^2}{\sigma_{x'}\sigma_{y'}}$$

$$B = \frac{F_e}{2\pi\sigma_x\sigma_y} = \frac{F_0}{2\pi\sigma_r^2} \frac{\sigma_r^2\sigma_{r'}^2}{\sigma_x\sigma_{x'}\sigma_y\sigma_{y'}}$$

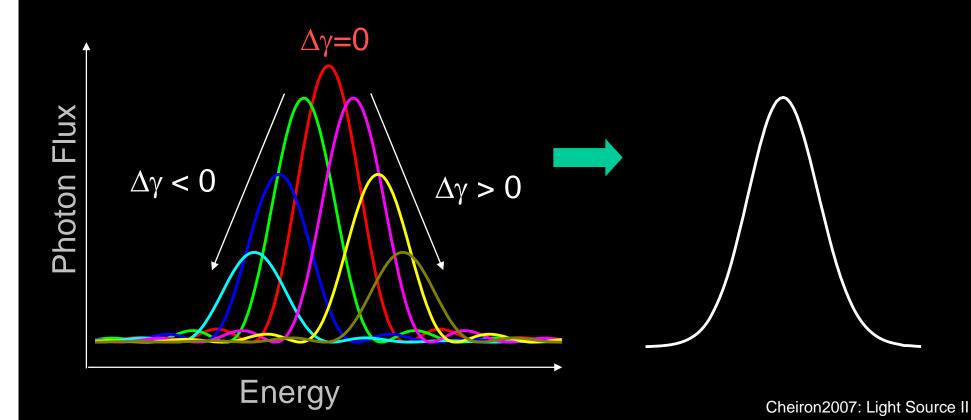
## Effects due to the Energy Spread

Electron with an offset of  $\Delta \gamma$ 



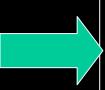
Energy shift of  $\omega_1$ 

$$\omega_1(\gamma) = \frac{4\pi c \gamma^2 / \lambda_u}{1 + K^2 / 2}$$

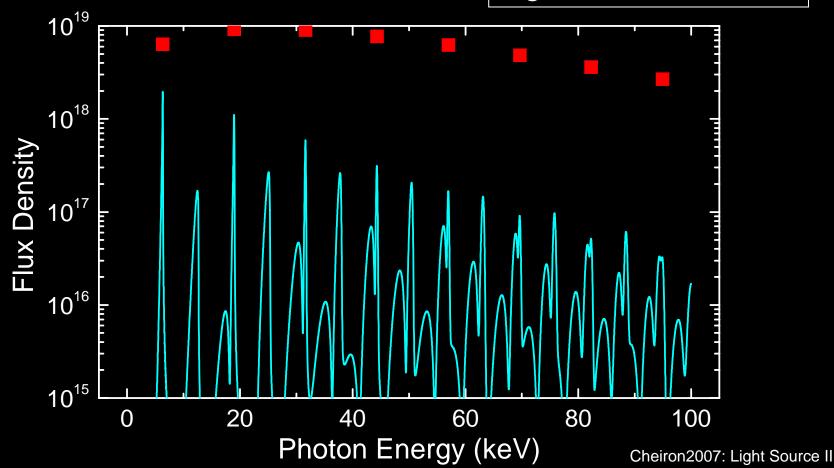


## Effects on the Higher Harmonics

Optical Emittance of UR:  $\lambda/4\pi$  Bandwidth of UR:  $\sim 1/nN$ 



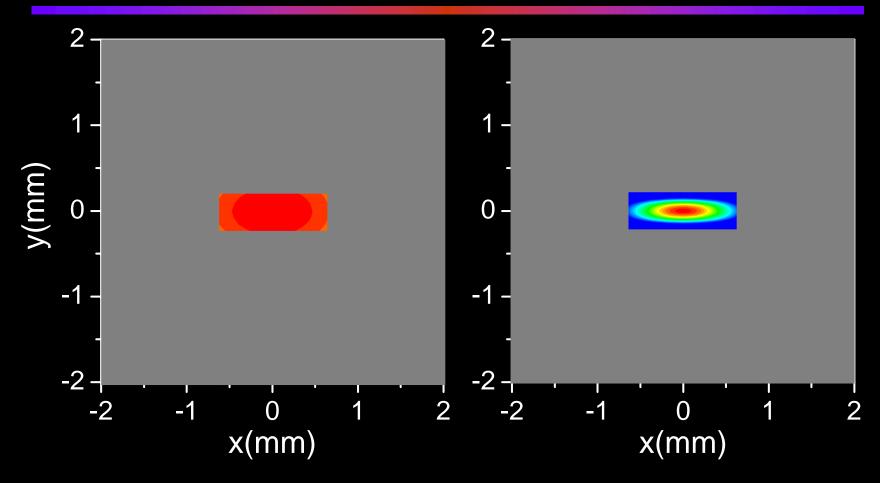
Effects due to the ebeam are larger for higher harmonics



## Heat Load on Optical Elements

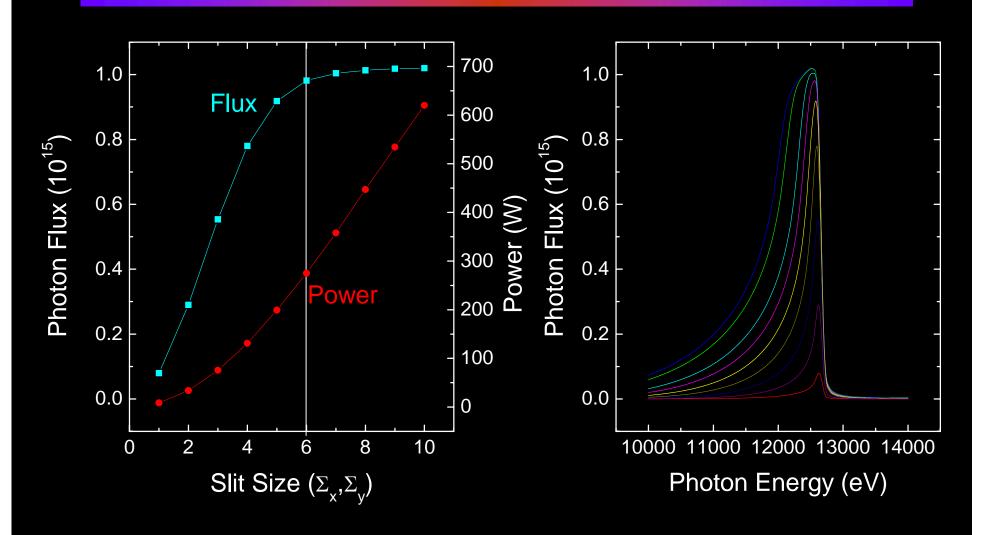
- SR emitted from the light source is processed by several optical elements before irradiation to the sample, such as the focusing mirror, monochromator.
- These elements can be easily damaged by the heat load brought by the SR.
- It is thus important to reduce the heat load as much as possible without sacrificing the flux, which is actually done by the XY slit at the front-end section.

## Spatial Profile of Power and Flux



The power profile is much broader than the flux. Extraction of SR with an appropriate slit significantly reduces the heat load.

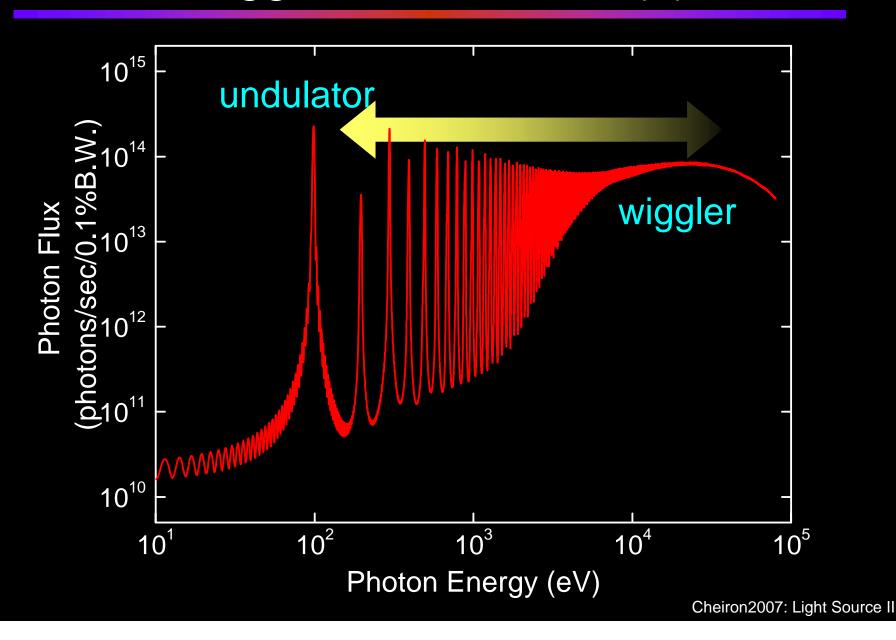
## Optimum Slit Size?



## Wiggler? Undulator? (1)

- Wigglers are identical to undulator from the point of view of magnetic circuit.
- It is generally said that the K value distinguishes between the two.
- However, this is not exactly correct.
- What we should take care is the region of photon energy to be utilized for application.

## Wiggler? Undulator? (2)



## Other Topics Not Addressed

- Quantitative descriptions of SR
- Light sources for circular polarization and schemes for fast helicity switching
  - helical undulator & elliptic wiggler
  - chicanes&choppers, kicker magnets
- Effects on the electron beam
  - natural focusing
  - beam-axis fluctuation due to COD variation
- R&Ds toward shorter magnetic period
  - superconducting undulators
  - cryogenic permanent magnet undulators
- Coherent SR for intense THz light
- Undulators for SASE-based X-ray FEL

#### **Announcements**

If you have your own PC, please download "SPECTRA" from the Web site

http://radiant.harima.riken.go.jp/spectra/index.html

for the lecture on Thu. 16:20~.

Thank you.